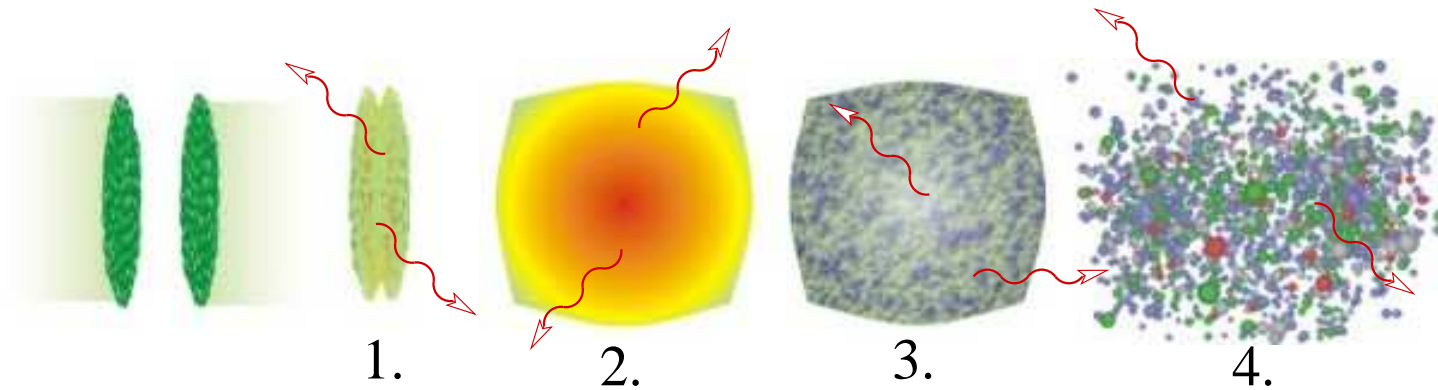


# Photons and Transport at NLO

with Jacopo Ghiglieri, Juhee Hong, Aleksi Kurkela, Egang Lu, Derek Teaney

- Photons: motivation and basics
- Convergence of Perturbation Theory
- When soft physics is light-cone physics
- When light-cone physics is thermodynamics
- NLO photon production: results, prospects

# Stages of a Heavy Ion Collision



1. Ions collide, making  $q, g$ , **photons** “primary”
2.  $q, g$  rescatter as QGP, make **photons** “thermal”?
3. Hadrons form, scatter, make **photons** “Hadronic”
4. Hadrons escape, some decay to **photons** “decay”

Photon re-interaction rare ( $\alpha_{\text{EM}} \ll 1$ ): direct info.

Thermal photons *may* act as a thermometer for QGP.

Production rate is interesting! Mostly for  $E > 2$  GeV, several  $T$ .

# How Photons Get Made

Since  $\alpha_{\text{EM}} \ll 1$ , work to lowest order in  $\alpha_{\text{EM}}$ :

- assume photon production *Poissonian* Find single-photon production
- neglect back-reaction on system cooling insignificant...

Single-photon production at  $\mathcal{O}(\alpha_{\text{EM}})$

$$2k^0 \frac{d\text{Prob}}{d^3k} = \sum_X \text{Tr} \rho U^\dagger(t) |X, \gamma(k)\rangle \langle X, \gamma(k)| U(t)$$

$U(t)$  time evolution operator,  $\rho$  density matrix.

Expand  $U(t)$  in EM interaction picture:

$$U(t) = 1 - i \int^t dt' \int d^3x e A^\mu(x, t') J_\mu(x, t') + \mathcal{O}(e^2)$$

$A^\mu$  produces the photon. Get

$$\frac{d\text{Prob}}{d^3k} = \frac{e^2}{2k^0} \int d^4Y d^4Z e^{-iK \cdot (Y-Z)} \sum_X \text{Tr } \rho J^\mu(Y) |X\rangle \langle X| J_\mu(Z)$$

And  $\sum_X |X\rangle \langle X| = \mathbf{1}$ . Assume slow-varying,  
near-equilibrium:  $\int d^4Z \rightarrow Vt$ : Get rate per 4-volume:

$$\frac{d\Gamma}{d^3k} = \frac{e^2}{2k^0} G^<(K), \quad G^<(K) \equiv \int d^4Y e^{-iK \cdot Y} \langle J^\mu(Y) J_\mu(0) \rangle_\rho$$

Success of Hydro – but not true at early times ....

# Computational Approaches

No first-principles, *nonperturbative* tool for  $\langle J^\mu J_\mu \rangle(K)$ . Only

- Lattice techniques (uncontrolled analytic continuation. Avoid)
- Weak-coupling techniques (uncontrolled extrapolation from  $\alpha_s < 0.1$ . Pursue)
- Strong-coupling  $\mathcal{N}=4$  SYM (uncontrolled relation to QCD. Avoid)

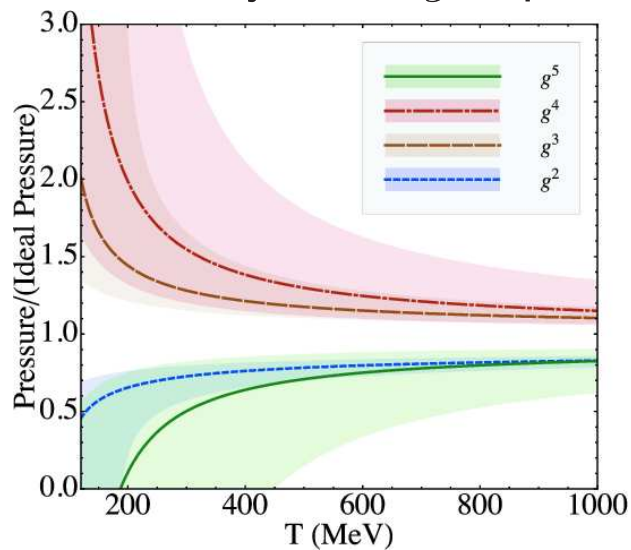
How bad is weak coupling?

- It fails at  $T \sim \text{few } T_C$  ?
- It fails at  $T \sim 10^6 T_C$  ?
- It fails at all temperatures?      **Truth: some of each!**

# Lessons from the Pressure

Divide degrees of freedom in 2 groups: *Hard and Soft*

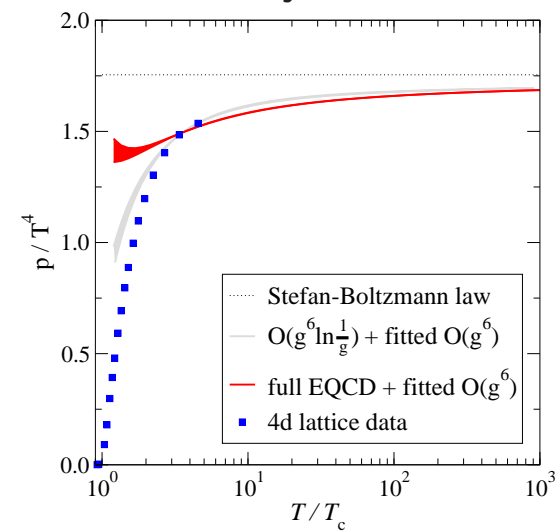
Naive order-by-order  $g$ -expansion



Converges if  $T > 10^6 T_C$

Arnold-Zhai, Braaten-Nieto, etc

Integrate out *Hard*, solve *Soft* nonperturbatively: *3-D theory!*



Works down to  $T = 2T_C$ .

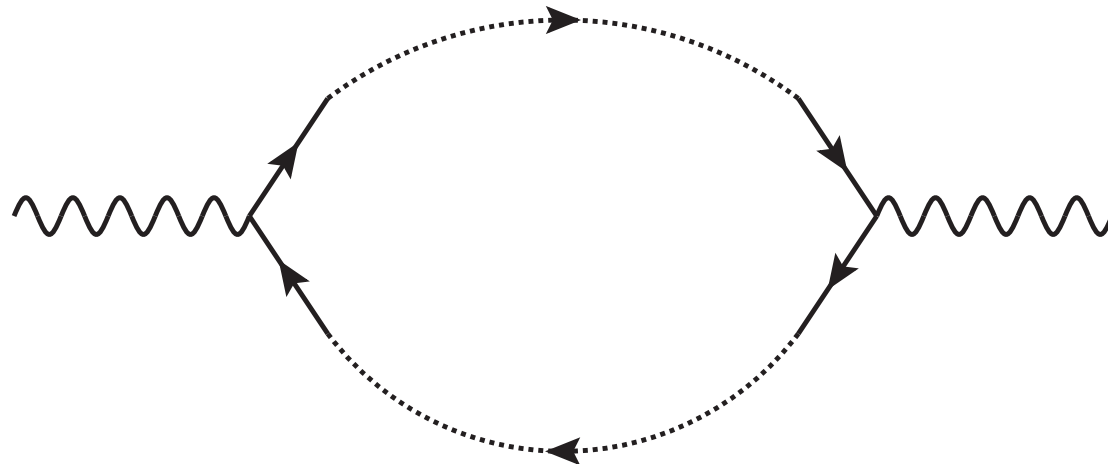
Kajantie et al, etc

Hard physics is perturbative. There is hope!

# Perturbative treatment

We want  $G^<(K) \equiv \int d^4Y e^{-iK \cdot Y} \langle J^\mu(Y) J_\mu(0) \rangle_\rho$

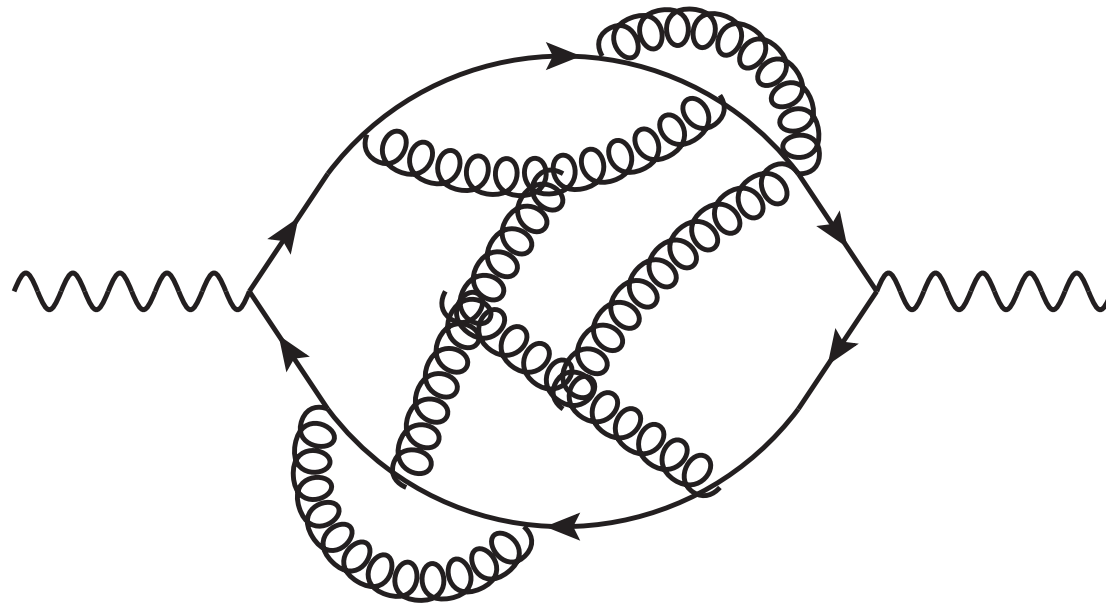
$J^\mu = \bar{\psi} \gamma^\mu \psi$ . Correlator of two quarks. Something like



# Perturbative treatment

We want  $G^<(K) \equiv \int d^4Y e^{-iK \cdot Y} \langle J^\mu(Y) J_\mu(0) \rangle_\rho$

$J^\mu = \bar{\psi} \gamma^\mu \psi$ . Correlator of two quarks. In general,



Worse: dynamics *complex*, no nice effective 3D theory!

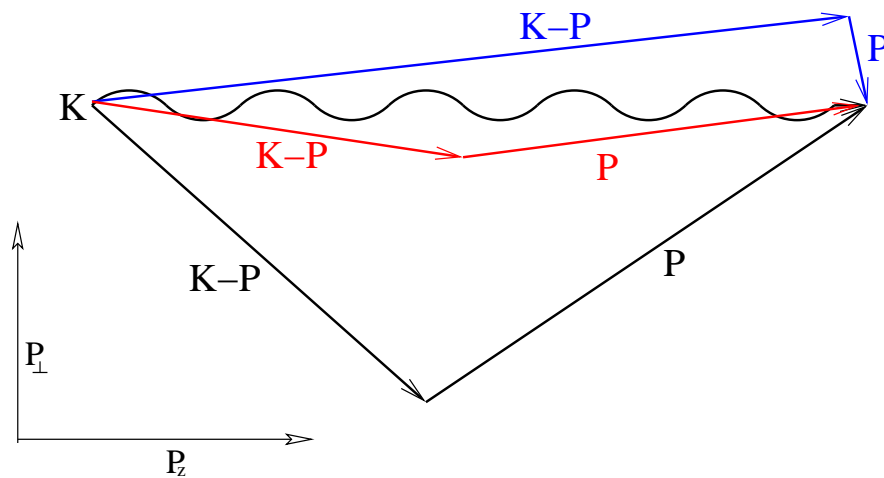


# Start with Kinematics

$$\overline{\mathcal{M}} \quad \mathcal{M}$$

$$\gamma \text{ produc: } \sum_{\psi_f} \langle \psi_i | A^\mu \bar{\psi} \gamma_\mu \psi | \psi_f \rangle \langle \psi_f | A^\nu \bar{\psi} \gamma_\nu \psi | \psi_i \rangle$$

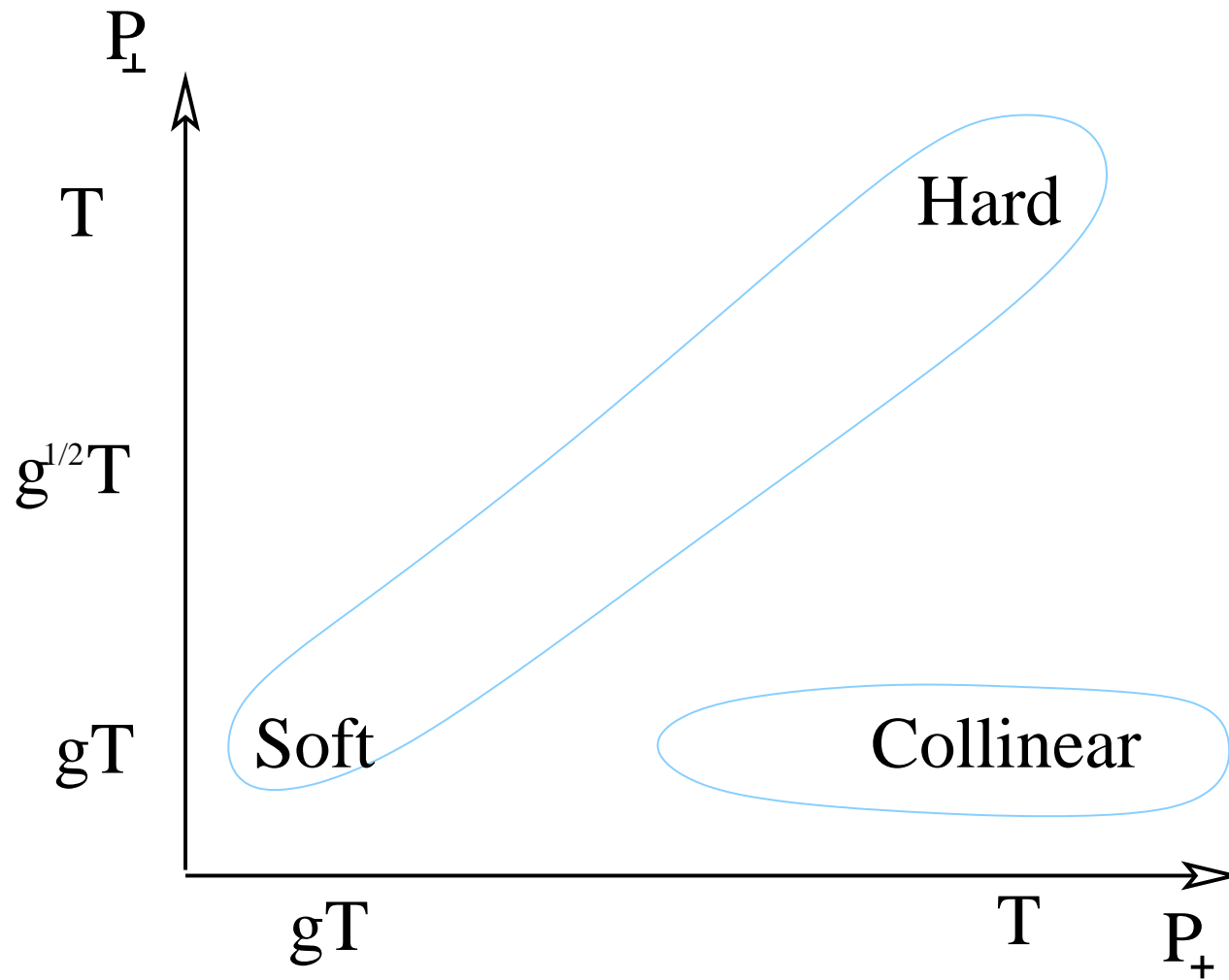
In  $\mathcal{M}$ ,  $\psi, \bar{\psi}$  momenta  $p, k - p$  must add to  $k$  of photon:



Black: way off-shell,  
but big phase space  
Blue: less phase sp,  
but soft enhancement  
Red: both can be  
almost on-shell.

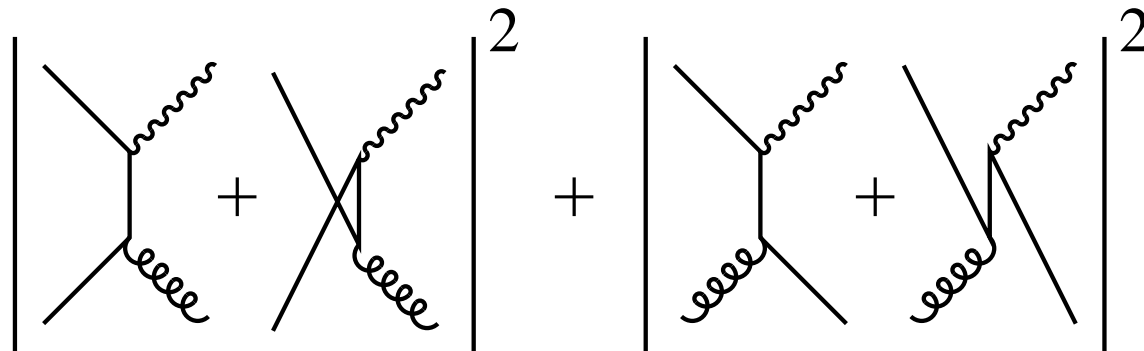
Call these regions Hard, Soft, and Collinear.

# The $P_{\perp}, P_{+}$ plane:



## Hard case

If all momentum components (transverse and longitudinal) are large, physics is simple: short distance-and-time correlators, PQCD works. Loop corrections are  $\mathcal{O}(g^2)$  and should get large around  $T \sim 2T_C$ .



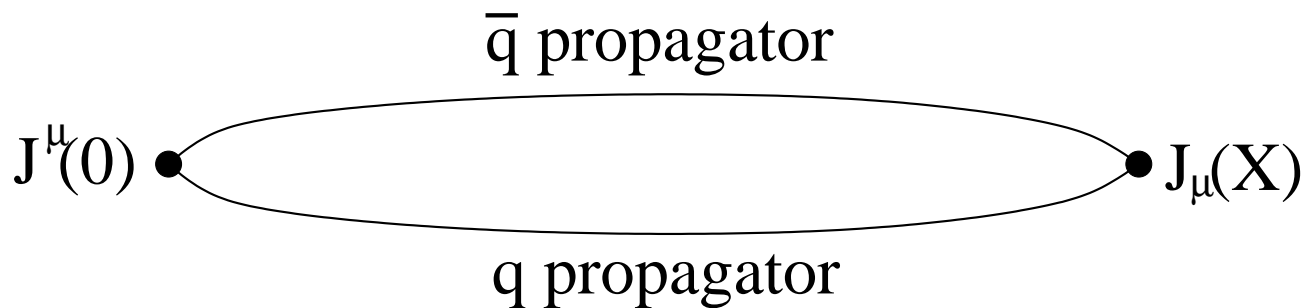
The challenge is the other two regions, where Pert. Thy. need not work as well.

# Momentum-space vs Coordinate space

Momentum  $K$  lightlike  $\not\leftrightarrow$  lightlike  $X$ -separation:

$$f(K) = \int d^4 X e^{-iK \cdot X} f(X) \quad \text{involves **all** } X$$

But  $q, \bar{q}$  start and end at same place:

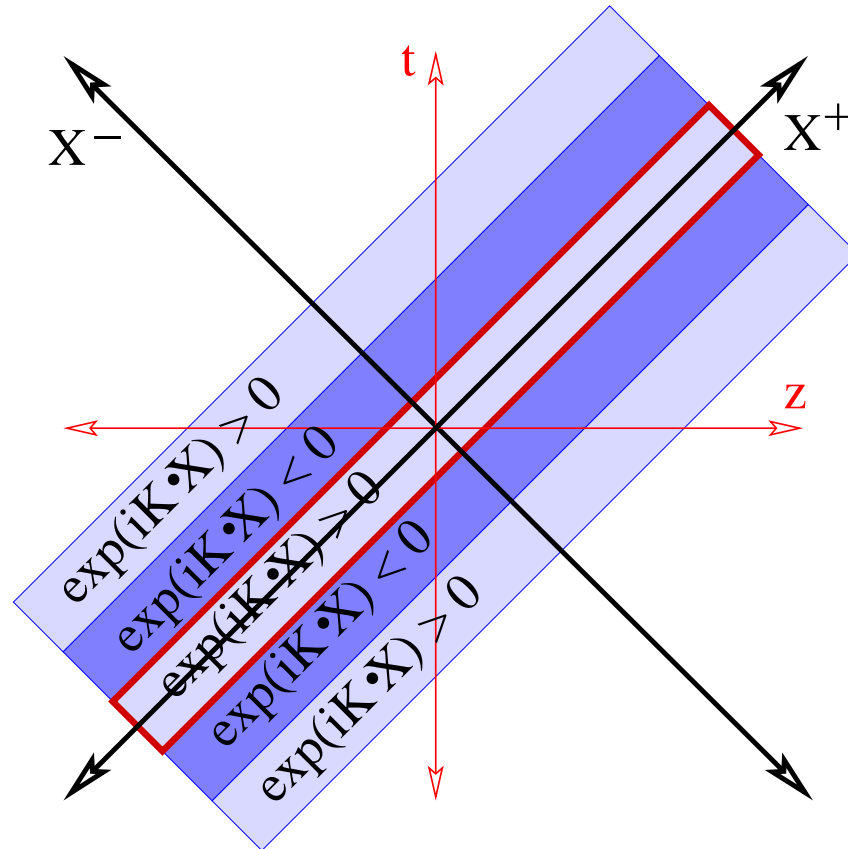


$X$  determined by Fourier properties of  $P$  and  $K - P$ .

$P$  small (or  $p_\perp$  small):  $X$  (or  $x_\perp$ ) large.

$K$  big, but  $X$  in  $\int d^4X \exp(-iK \cdot X)$  also big. How?

Need phase  
 $\exp(-iK \cdot X)$   
 small. Occurs in  
 narrow region.  
 Write  $t, z$  as  
 $X^- = (t - z),$   
 $X^+ = (t + z)/2.$



Since  $-K \cdot X = K^+ X^- + K^- X^+$ ,  $K^+$  big,  
 contribution is from region  $X^- \simeq 0$  (Light Cone)

# Lightcone correlators are Simple!

$x^- = 0$  ( $x = t$ ) is “Lightcone” of photon

Separation lightlike if  $x_{\perp} = 0$ , spacelike if  $x_{\perp} \neq 0$ .

Causality  $\rightarrow$  only *pre-existing* correlators.

Unequal times usually means *Complicated Dynamics*.

Now ~~Complicated dynamics~~ Simple Thermodynamics!

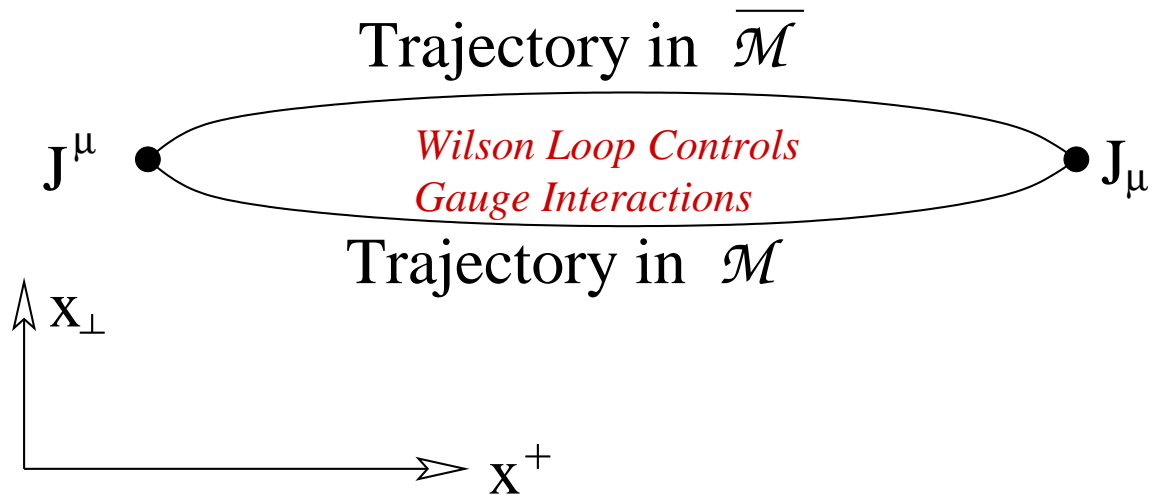
- Energy-dependent: Just Thermal Masses!
- Energy-independent: Classical (3-D theory) correlators!

# Collinear case

Collinear  $\Rightarrow$  almost on-shell  $\Rightarrow$  large  $x$  separation

$$x^- \ll x_\perp \ll x^+ \quad (1/T \ll 1/gT \ll 1/g^2T)$$

Consider *spacetime trajectory* of  $q, \bar{q}$ :



Need  $x_\perp$ -separated Wilson loop.

Spacetime picture pioneered by B. Zakharov, hep-ph/9607440,9807540

# Nontrivial analysis B. Zakharov, BDMPS, AMY

$$\begin{aligned}
 \frac{dN_\gamma}{d^3\mathbf{k}d^4x} &= \frac{\alpha_{\text{EM}}}{\pi^2 k} \int_{-k/2}^{\infty} \frac{dp^+}{2\pi} n_f(k+p) [1-n_f(p)] \frac{p^2 + (p+k)^2}{2[p(p+k)]^2} \\
 &\quad \times \lim_{\mathbf{x}_\perp \rightarrow 0} 2 \text{Re} \partial_{\mathbf{x}_\perp} \mathbf{f}(x_\perp) \\
 2\nabla_\perp \delta^2(x_\perp) &= \left[ \underset{\text{J-operator}}{\mathcal{C}(x_\perp)} + \frac{ik}{2p^+(k+p^+)} \underset{x_\perp\text{-diffusion}}{(m_\infty^2 + \nabla_{x_\perp}^2)} \right] \mathbf{f}(x_\perp) \\
 &\quad \underset{\text{Wilson-line}}{\mathcal{C}(x_\perp)}
 \end{aligned}$$

$\mathbf{f}(x_\perp)$ : density matrix  $|\psi_{P+K}\rangle\langle\gamma_K\psi_P|$  or  $|\psi_P\bar{\psi}_{K-P}\rangle\langle\gamma_K|$

Eikonal evolution (Evolution in  $x^+$ ) –  $x_\perp$  diffusion, AND

Wilson-loop interaction with medium  $\mathcal{C}(x_\perp)$ .



# $\mathcal{C}(x_\perp)$ is Euclidean!

$\mathcal{C}(x_\perp)$ : Wilson loop with space-separated lightlike lines. All points at spacelike or lightlike separation.

Soft contribution is Euclidean!! [S. Caron-Huot, 0811.1603](#)

Calculate it with *simple* perturbation theory (EQCD)

Calculate it on the lattice?!

NLO corrections to  $\mathcal{C}(x_\perp)$  computed. NNLO would be nonperturbative; possible via lattice [Rummukainen, next talk](#)

# How Things Get Euclidean S. Caron-Huot

Consider correlator  $G^<(x^0, \mathbf{x})$  with  $x^z > |x^0|$ . Fourier representation

$$G^<(x^0, \mathbf{x}) = \int d\omega \int dp_z d^2 p_\perp e^{i(x^z p^z + \mathbf{x}_\perp \cdot \mathbf{p}_\perp - \omega x^0)} G^<(\omega, p_z, p_\perp)$$

Use  $G^<(\omega, \mathbf{p}) = n_b(\omega)(G_R(\omega, \mathbf{p}) - G_A(\omega, \mathbf{p}))$  and define  $\tilde{p}^z = p^z - (t/x^z)\omega$ :

$$G^< = \int d\omega \int d\tilde{p}^z d^2 p_\perp e^{i(x^z \tilde{p}^z + \mathbf{x}_\perp \cdot \mathbf{p}_\perp)} n_b(\omega) \left( G_R(\omega, \tilde{p}^z + \omega \frac{x^0}{x^z}, \mathbf{p}_\perp) - G_A \right)$$

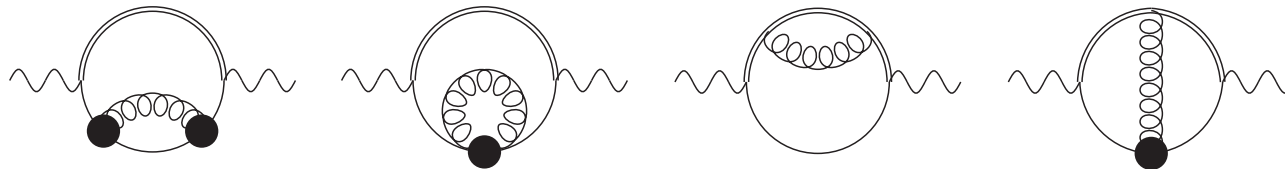
Perform  $\omega$  integral: upper half-plane for  $G_R$ , lower for  $G_A$ , pick up poles from  $n_b$ :

$$G^<(x^0, \mathbf{x}) = T \sum_{\omega_n = 2\pi n T} \int dp^z d^2 p_\perp e^{i\mathbf{p} \cdot \mathbf{x}} G_E(\omega_n, p_z + i\omega_n(x^0/x^z), p_\perp)$$

Large separations:  $n \neq 0$  exponentially small.  $n = 0$  contrib. is  $x^0$  independent!

# Soft momenta

Start with brute force: do the diagrams



Cut hard line:  $p^- \simeq 0$ , hard-line approx.  $p^+$  independent.

Remaining integrals (using KMS) ( $P, Q$  are resp. soft quark, gluon momenta)

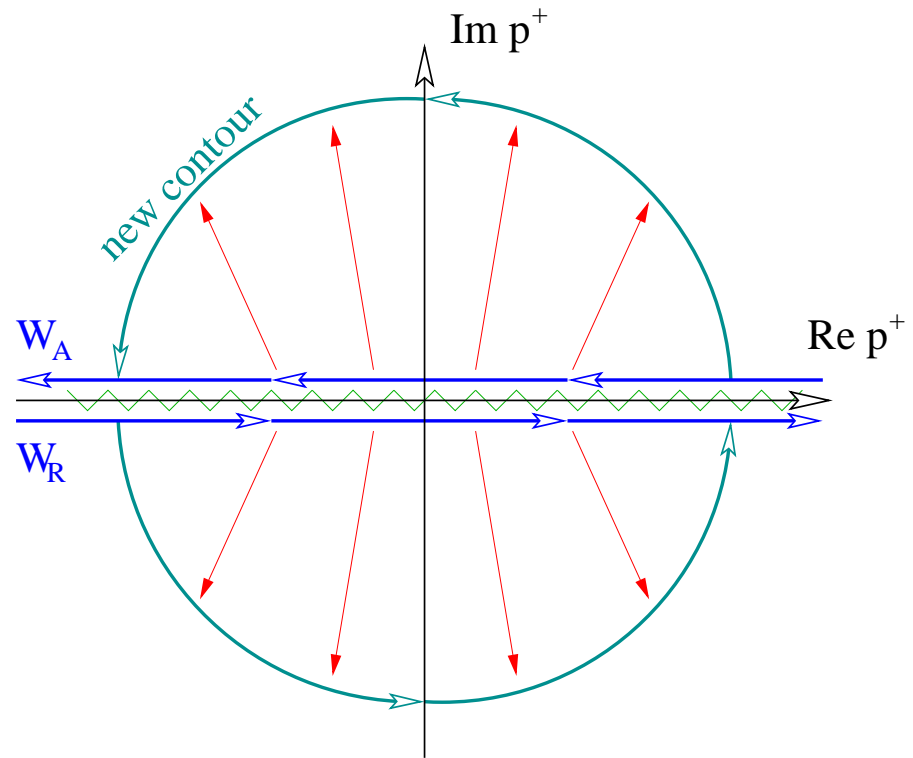
$$\int_{\sim gT} d^2 p_{\perp} dp^+ \int_{\sim gT} d^4 Q n_b(k^0) (G_R - G_A)$$

$G_R$ : retarded function of sum of all 4 diagrams' guts.

Momentum  $p^+$  is **null**. Any  $R/A$  function is analytic in upper/lower half plane for time-like or **null**  $p$ -variable.

Analytically continue in  $p^+!!$

Deform  $p^+$  contour  
into complex plane



Now  $p^+ \gg p_\perp, Q$ . (On mass-shell) Expand in  $p^+ \gg p_\perp, Q$

$$G_R[4 \text{ diagrams}] = C_0(p^+)^0 + C_1(p^+)^{-1} + \dots$$

$C_0$  is on-shell width, gives linear in  $p^+$  divergence.

$C_1$  is on-shell dispersion correction,  $dp^+/p^+$  gives const.

**Huh?** Continuation possible because  $J^\mu$  light-cone separated. And light-cone correlators are simple!

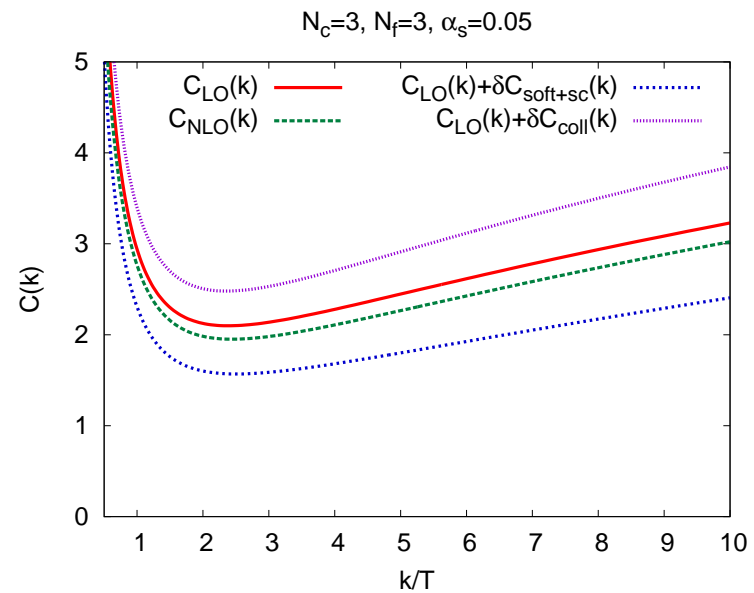
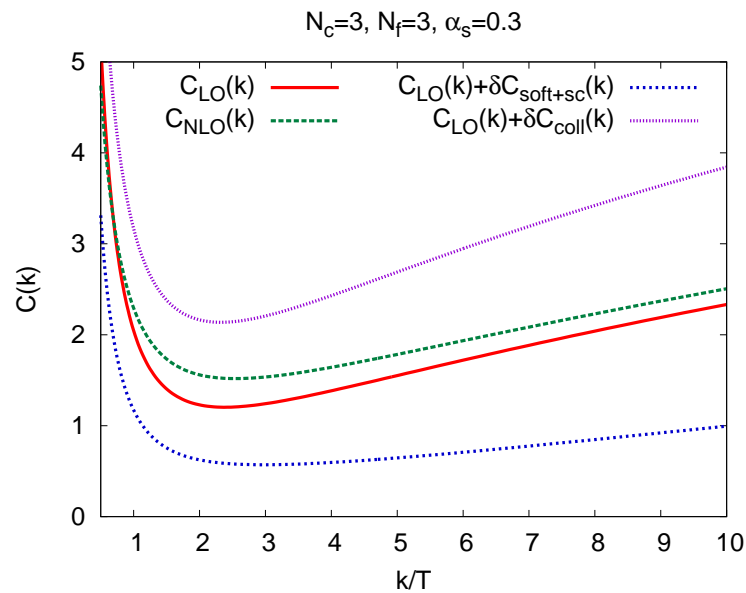
- $C_0$  term: Exactly the limit of collinear calculation when one quark momentum gets small. Already included.
- $C_1$  term: real dispersion-correction. Really simple:

$$\gamma\text{-rate} \propto \int \frac{d^2 p_\perp}{(2\pi)^2} \frac{m_\infty^2}{p_\perp^2 + m_\infty^2}$$

where  $m_\infty^2$  is dispersion correction. Has leading-order piece (hard modes) and subleading piece (dispersion correction of soft modes). *both are known.*

Remaining region—similar story. Null-separation physics, all condensates.

Summing it up: two corrections



Upward correction: more scattering at NLO.

Downward correction: fewer soft gluons, less dispersion corr.

Numerical conspiracy: effects nearly cancel **[Accidental!]**

# Main lesson

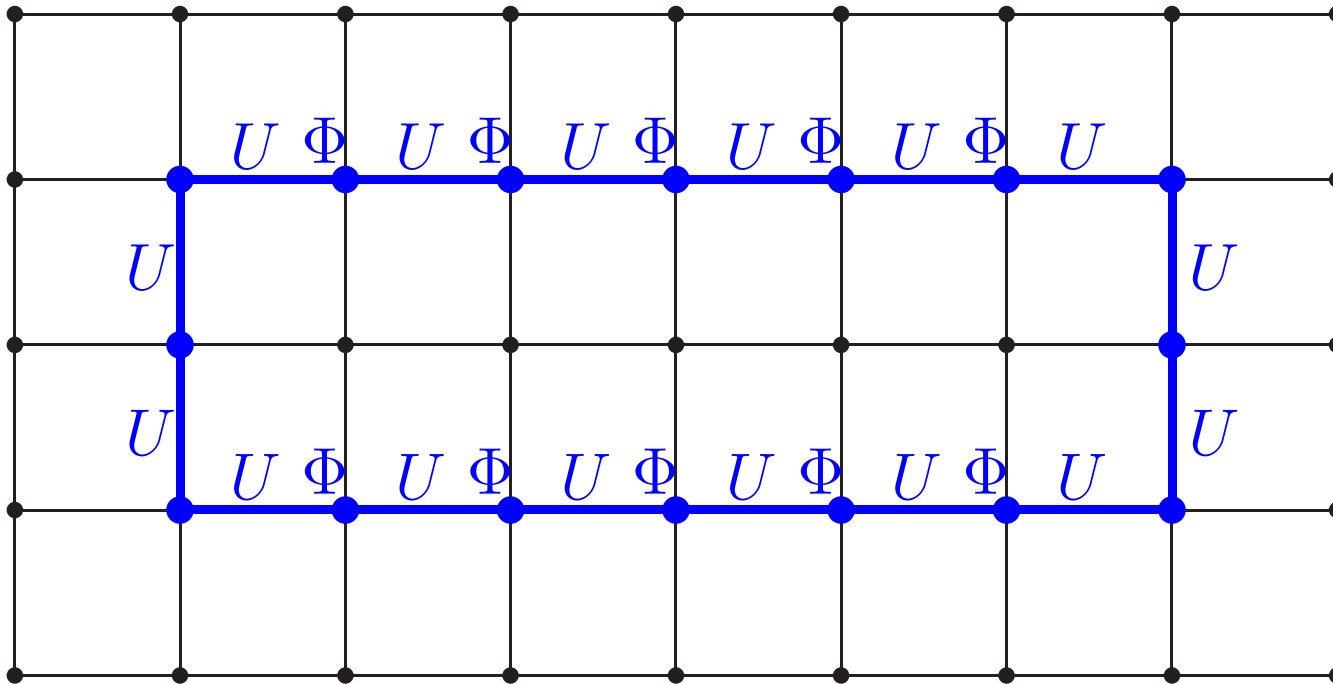
All the sticky IR physics shows up in a few condensates.  
Some are dispersion corrections – physically simple.  
Some are Euclidean – get directly on the lattice.

**Bad news:**  $\mathcal{O}(g)$  corrections big even for  $\alpha_s = 0.1$  or  $1000 T_c$ . *Sort of expected that.*

**Good news:** A few condensates. Determine them nonperturbatively, maybe get down to few  $T_c$ ?

Get them on the lattice?

## $\mathcal{C}(x_\perp)$ on the lattice



Short side:  $x_\perp$  Wilson line  $\exp \int i A_\perp \cdot x_\perp \Rightarrow U_\perp U_\perp \dots$

Long side:  $x^+$  Wilson line  $\exp \int i (A^z + A^0) dz \Rightarrow U_z e^{a\Phi} U_z e^{a\Phi} U_z \dots$

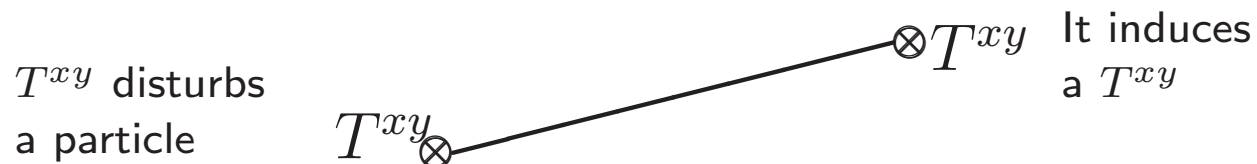
The latter is a new beast. Lattice renormalization properties?

Under investigation.

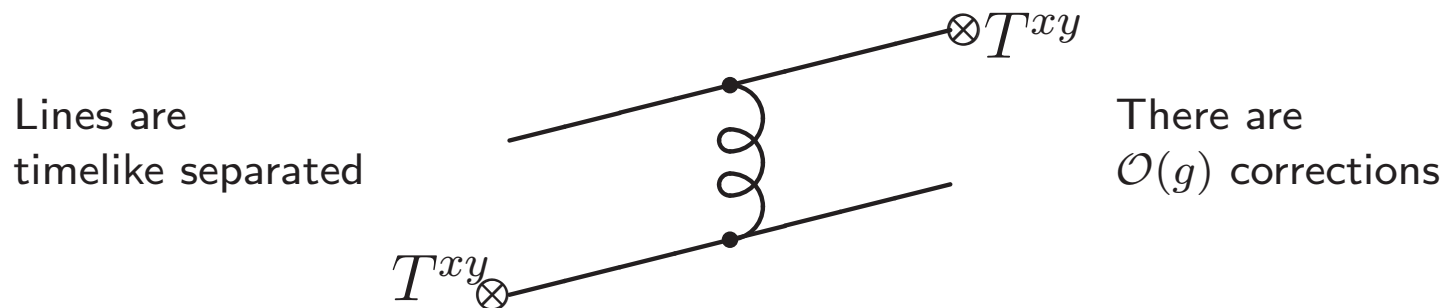


# Other transport coefficients?

We want Baryon Diffusion  $D$  and (especially) shear  $\eta$ !  
Both controlled by high-energy  $E =$  several  $T$  particles  
Lightlike correlators should again dominate:



NLO effects arise along particle's lightlike trajectory.  
Problem: transfer of stress to someone else



# Conclusions

- Photon production is worth computing
- “Enhanced” Pert. calculation – few  $T_C$ ??
- NLO corrections to transport are *large but simple*
- Need a few correlators at lightlike-separated points
- Most can be extracted from the lattice
- Shear and diffusion will be harder. Stay tuned